

Marseille Abstract

In this talk, I will present UV continuum properties from the EPOCHS sample of 1011 high-redshift galaxies spanning $6.5 < z < 13$ across 179 sq arcmin of public and PEARLS GTO NIRCcam imaging from JWST Cycle 1. I will show the bias corrected beta-MUV relation, from which I find that the MUV=-19 galaxy population becomes extremely blue at $z > 11$. The EPOCHS sample indeed contains a subsample of 68 robust ultra-blue galaxies with $\beta < -2.8$ with high Lyman continuum escape fractions which may be important contributors to the reionization of the Universe at early times. From the EPOCHS beta-stellar mass scaling relation, I find a large population of faint, low mass, red galaxies at $6.5 < z < 11$ and a non-detection of these sources at $z > 11$. The steep slope of beta-stellar mass at $z > 11$ implies that at early times type II SNe are the major dust producers and that low mass $< 10^8$ solar galaxies with smaller gravitational potential wells lose this dust in outflows induced by SNe feedback. Carbon-rich Wolf-Rayet (WR) binaries may well be the culprits of the dust produced in low mass galaxies at $z=9$ since dust production in the winds of AGB stars takes longer than the rapid build up implied by the observed red beta slopes. To conclude I will talk about how the improved sensitivity of ELT/MICADO compared to JWST/NIRCcam will improve stellar mass measurements at $z > 6.5$ as a result of the mitigation of outshining via resolved SED fitting. This will inevitably lead to a greater understanding of the sources of dust production in the early Universe.